

Chapter 3

Rotational motion and the hydrogen atom

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Exercises

3.1

$$E = m_l^2 (\hbar^2/2I) \text{ [eqn 3.6]; } I = m_H R^2$$

$$E = m_l^2 (\hbar^2/2m_H R^2)$$

$$m_H = 1.674 \times 10^{-27} \text{ kg, } R = 160 \text{ pm; } \hbar^2/2m_H R^2 = 1.30 \times 10^{-22} \text{ J}$$

Hence,

$$E = \underline{(1.30 \times 10^{-22} \text{ J})m_l^2}$$

3.2 Using the energy levels from Exercise 3.1, we obtain

$$\Delta E = (1.30 \times 10^{-22} \text{ J})(1 - 0) = 1.30 \times 10^{-22} \text{ J}$$

$$\lambda = hc/\Delta E = 1.53 \times 10^{-3} \text{ m} = \underline{1.53 \text{ mm}}$$

This wavelength corresponds to microwave radiation.

Exercise: Calculate the effect of deuteration on E and $\lambda(1 \leftarrow 0)$.

$$\mathbf{3.3} \quad x = r \cos \phi, \quad y = r \sin \phi, \quad r = (x^2 + y^2)^{1/2}$$

$$\begin{aligned}
 l_z &= \frac{\hbar}{i} \left(x \frac{\partial}{\partial y} - y \frac{\partial}{\partial x} \right) \\
 &= \frac{\hbar}{i} \left(x \left\{ \frac{\partial r}{\partial y} \frac{\partial}{\partial r} + \frac{\partial \phi}{\partial y} \frac{\partial}{\partial \phi} \right\} - y \left\{ \frac{\partial r}{\partial x} \frac{\partial}{\partial r} + \frac{\partial \phi}{\partial x} \frac{\partial}{\partial \phi} \right\} \right) \\
 &= \frac{\hbar}{i} \left(\left\{ \frac{xy}{r} \frac{\partial}{\partial r} + x \frac{\partial \phi}{\partial y} \frac{\partial}{\partial \phi} \right\} - \left\{ y \frac{x}{r} \frac{\partial}{\partial r} + y \frac{\partial \phi}{\partial x} \frac{\partial}{\partial \phi} \right\} \right) \\
 &= \frac{\hbar}{i} \left(x \frac{\partial \phi}{\partial y} - y \frac{\partial \phi}{\partial x} \right) \frac{\partial}{\partial \phi} \\
 &= \frac{\hbar}{i} \left(\frac{x}{\partial y / \partial \phi} - \frac{y}{\partial x / \partial \phi} \right) \frac{\partial}{\partial \phi} \\
 &= \frac{\hbar}{i} \left(\frac{r \cos \phi}{r / \cos \phi} + \frac{r \sin \phi}{r / \sin \phi} \right) \frac{\partial}{\partial \phi} = \frac{\hbar}{i} \frac{\partial}{\partial \phi}
 \end{aligned}$$

3.4

$$\begin{aligned}
 \int_0^{2\pi} \phi_{m_l}^* \phi_{m_l'} d\phi &= (1/2\pi) \int_0^{2\pi} e^{i(m_l - m_l')\phi} d\phi \\
 &= (1/2\pi) \left\{ \frac{e^{2i(m_l - m_l')\pi} - 1}{(m_l - m_l')i} \right\} = 0 \quad \text{if } m_l' \neq m_l \\
 &\quad [e^{2in\pi} = 1, n \text{ an integer}]
 \end{aligned}$$

(Note that when $m_l' = m_l$ the integral has the value 2π .)

Exercise: Normalize the wavefunction $e^{i\phi} \cos \beta + e^{-i\phi} \sin \beta$, and find an orthogonal linear combination of $e^{+i\phi}$ and $e^{-i\phi}$.

3.5 The moment of inertia of a solid uniform disc of mass M and radius R is

$$I = \frac{1}{2} MR^2; \quad \text{hence } I = 2.5 \times 10^{-4} \text{ kg m}^2$$

Then

$$E = m_l^2 (\hbar^2 / 2I) = \underline{(2.2 \times 10^{-65} \text{ J}) m_l^2}$$

The rotation rate is 100 Hz. Hence $\omega = 2\pi\nu = 628 \text{ s}^{-1}$. The angular momentum is $I\omega = 0.16 \text{ kg m}^2 \text{ s}^{-1}$. If this is set equal to $|m_l|\hbar$ we require $|m_l| = 1.5 \times 10^{33}$. Since the disc rotates anticlockwise when seen from below, m_l is negative. Hence, $\underline{m_l = -1.5 \times 10^{33}}$.

Exercise: How much more energy is required to raise the disc into its next rotational state?

3.6 See Fig. 3.1. We have plotted

$$\text{re } \phi_{m_l} = (1/2\pi)^{1/2} \cos m_l \phi = \begin{cases} (1/2\pi)^{1/2} \cos 3\phi & \text{for } m_l = 3 \\ (1/2\pi)^{1/2} \cos 4\phi & \text{for } m_l = 4 \end{cases}$$

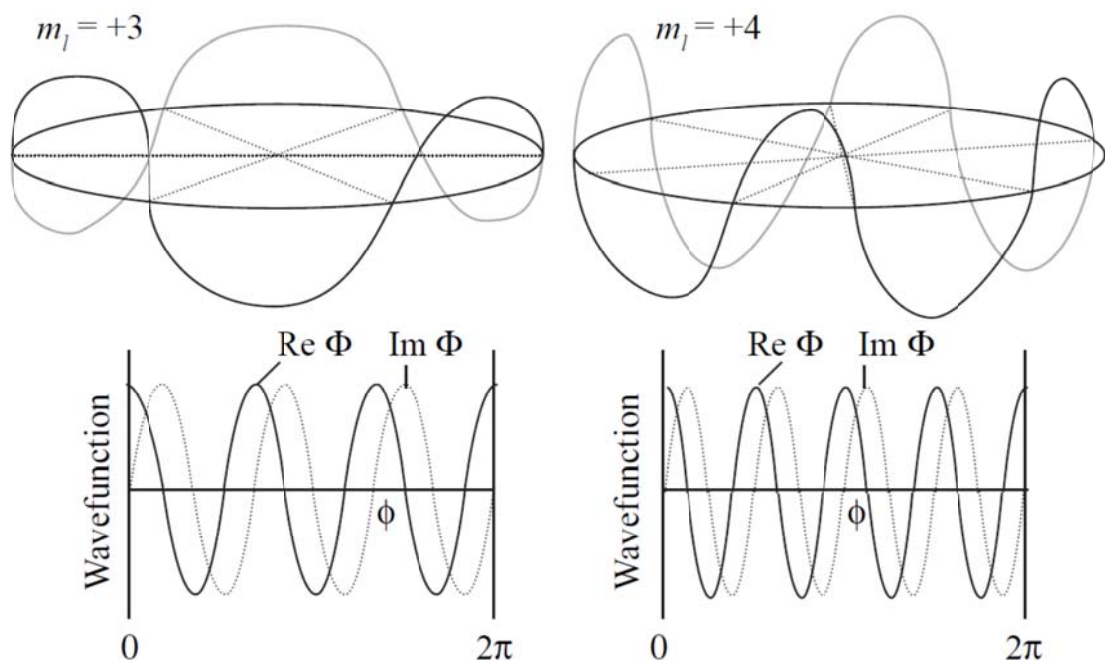


Figure 3.1: A representation of the amplitudes and phases of the wavefunction of a particle on a ring (red, real; green, imaginary).

Exercise: Superimpose the imaginary parts of ϕ_{m_l} on the diagrams. Draw $\text{Re } \phi$ for the superposition $e^{i\phi} \cos \beta + e^{-i\phi} \sin \beta$.

3.7 The Schrödinger equation is given in eqn 3.31:

$$\Lambda^2 \psi = -(2IE/\hbar^2) \psi; \quad \Lambda^2 = (1/\sin^2 \theta)(\partial^2/\partial \phi^2) + (1/\sin \theta)(\partial/\partial \theta) \sin \theta (\partial/\partial \theta)$$

Write $\psi = \Theta\Phi$; then with $\Theta' = d\Theta/d\theta$ and $\Phi' = d\Phi/d\phi$, etc.

$$(1/\sin^2 \theta)\Theta\Phi'' + (1/\sin \theta)\Phi(d/d\theta) \sin \theta \Theta' = -(2IE/\hbar^2)\Theta\Phi$$

$$\Phi''/\Phi + (1/\Theta) \sin \theta (d/d\theta) \sin \theta \Theta' = -(2IE/\hbar^2) \sin^2 \theta$$

Write $\Phi''/\Phi = -m_l^2$, a constant; then

$$(1/\Theta) \sin \theta (d/d\theta) \sin \theta \Theta' = m_l^2 - (2IE/\hbar^2) \sin^2 \theta$$

Because $(d/d\theta) \sin \theta \Theta' = \Theta' \cos \theta + \Theta'' \sin \theta$, this rearranges into

$$\Theta'' \sin^2 \theta + \Theta' \sin \theta \cos \theta = \{m_l^2 - (2IE/\hbar^2) \sin^2 \theta\} \Theta$$

Exercise: Identify this equation in M. Abramowitz and I.A. Stegun, *Handbook of mathematical functions*, and write down its solutions.

3.8 The Schrödinger equation is

$$\Lambda^2 \psi = -(2IE/\hbar^2) \psi \quad [\text{eqn 3.31}]$$

Write $\psi = \Theta(\theta)\Phi(\phi)$, $\Lambda^2 = \frac{1}{\sin^2 \theta} \frac{\partial^2}{\partial \phi^2} + \frac{1}{\sin \theta} \frac{\partial}{\partial \theta} \sin \theta \frac{\partial}{\partial \theta}$ [eqn 3.29], and $2IE/\hbar^2 = \varepsilon^2$;

then

$$\frac{1}{\sin^2 \theta} \Theta \frac{d^2 \Phi}{d\phi^2} + \frac{\Phi}{\sin \theta} \frac{d}{d\theta} \sin \theta \frac{d\Theta}{d\theta} = -\varepsilon^2 \Phi \Theta$$

Divide through by $\Theta\Phi$ and multiply through by $\sin^2 \theta$:

$$\frac{1}{\Phi} \frac{d^2 \Phi}{d\phi^2} + \frac{\sin \theta}{\Theta} \frac{d}{d\theta} \sin \theta \frac{d\Theta}{d\theta} + \varepsilon^2 \sin^2 \theta = 0$$

Write $\frac{1}{\Phi} \frac{d^2 \Phi}{d\phi^2} = -m_l^2$, so $\frac{d^2 \Phi}{d\phi^2} = -m_l^2 \Phi$ which implies that

$$\frac{\sin \theta}{\Theta} \frac{d}{d\theta} \sin \theta \frac{d\Theta}{d\theta} + \varepsilon^2 \sin^2 \theta = m_l^2$$

and hence that

$$\sin \theta \frac{d}{d\theta} \sin \theta \frac{d\Theta}{d\theta} + \varepsilon^2 \sin^2 \theta \Theta - m_l^2 \Theta = 0$$

and the equation is separable.

Exercise: Is the equation separable if $V(\theta, \phi) = af(\theta) + bg(\phi)$?

3.9 It is sufficient to show that the Y_{lm_l} satisfy $\Lambda^2 Y_{lm_l} = -l(l+1)Y_{lm_l}$ [eqn 3.33].

$$Y_{11} = -\frac{1}{2} (3/2\pi)^{1/2} \sin \theta e^{i\phi}$$

$$\begin{aligned}
 \Lambda^2 \sin \theta e^{i\phi} &= (1/\sin^2 \theta)(\partial^2/\partial\phi^2) \sin \theta e^{i\phi} + (1/\sin \theta)(\partial/\partial\theta) \sin \theta (\partial/\partial\theta) \sin \theta e^{i\phi} \\
 &= -(1/\sin \theta)e^{i\phi} + (1/\sin \theta)e^{i\phi}(d/d\theta) \sin \theta \cos \theta \\
 &= -(1/\sin \theta)e^{i\phi} + (1/\sin \theta)e^{i\phi}(\cos^2 \theta - \sin^2 \theta) \\
 &= -(1/\sin \theta)e^{i\phi} + (1/\sin \theta)e^{i\phi}(1 - 2 \sin^2 \theta) = -2 \sin \theta e^{i\phi}
 \end{aligned}$$

Hence, $\Lambda^2 Y_{11} = -2Y_{11}$, in accord with $l = 1$.

$$Y_{20} = \frac{1}{4} (5/\pi)^{1/2} (3 \cos^2 \theta - 1)$$

$$\begin{aligned}
 \Lambda^2 (3 \cos^2 \theta - 1) &= (1/\sin \theta)(d/d\theta) \sin \theta (d/d\theta)(3 \cos^2 \theta - 1) \\
 &= -6(1/\sin \theta)(d/d\theta) \sin^2 \theta \cos \theta \\
 &= -6(1/\sin \theta) \{2 \sin \theta \cos^2 \theta - \sin^3 \theta\} \\
 &= -6 \{2 \cos^2 \theta - \sin^2 \theta\} = -6(3 \cos^2 \theta - 1)
 \end{aligned}$$

Hence, $\Lambda^2 Y_{20} = -6Y_{20}$, in accord with $l = 2$.

3.10

$$\begin{aligned}
 \int |Y_{11}|^2 d\tau &= \frac{1}{4} (3/2\pi) \int_0^\pi \sin^2 \theta \sin \theta d\theta \int_0^{2\pi} d\phi \\
 &= \frac{3}{4} \int_{-1}^1 (1-x^2) dx \quad [x = \cos \theta] = 1
 \end{aligned}$$

$$\begin{aligned}
 \int |Y_{20}|^2 d\tau &= \frac{1}{16} (5/\pi) \int_0^\pi (3 \cos^2 \theta - 1)^2 \sin \theta d\theta \int_0^{2\pi} d\phi \\
 &= \frac{5}{8} \int_{-1}^1 (3x^2 - 1)^2 dx = 1
 \end{aligned}$$

$$\int Y_{11}^* Y_{20} d\tau \propto \int_0^{2\pi} e^{-i\phi} d\phi = 0$$

Exercise: Repeat the calculation for Y_{21} and Y_{31} .

3.11 From eqn 3.44, $E = J(J + 1)(\hbar^2/2I) = (1.30 \times 10^{-22} \text{ J})J(J + 1)$ [Exercise 3.1]. Draw up the following Table, using degeneracy $g_J = 2J + 1$:

J	$E/(10^{-22} \text{ J})$	g_J
0	0	1
1	2.60	3
2	7.80	5

3.12 Using the energies in Exercise 3.11, we find

$$\Delta E(1 - 0) = 2.60 \times 10^{-22} \text{ J}$$

$$\lambda(1 - 0) = hc/\Delta E(1 - 0) = 7.64 \times 10^{-4} \text{ m} = \underline{0.764 \text{ mm}} \text{ (far infrared)}$$

Exercise: Calculate the same quantities for the deuterated species.

3.13 See Fig. 3.2. From Problem 3.11, when $l = 3$ and $m_l = 0$, $\theta = 90^\circ$, and the angular momentum vector lies in the equatorial plane; therefore $|Y|^2$ will have maxima on the z -axis, as seen in Fig. 3.2. We also see from Problem 3.11 that as $|m_l|$ increases, the deviation of θ from 90° increases; as the projection of the angular momentum vector on the z -axis increases, $|Y|^2$ becomes larger in the equatorial plane.

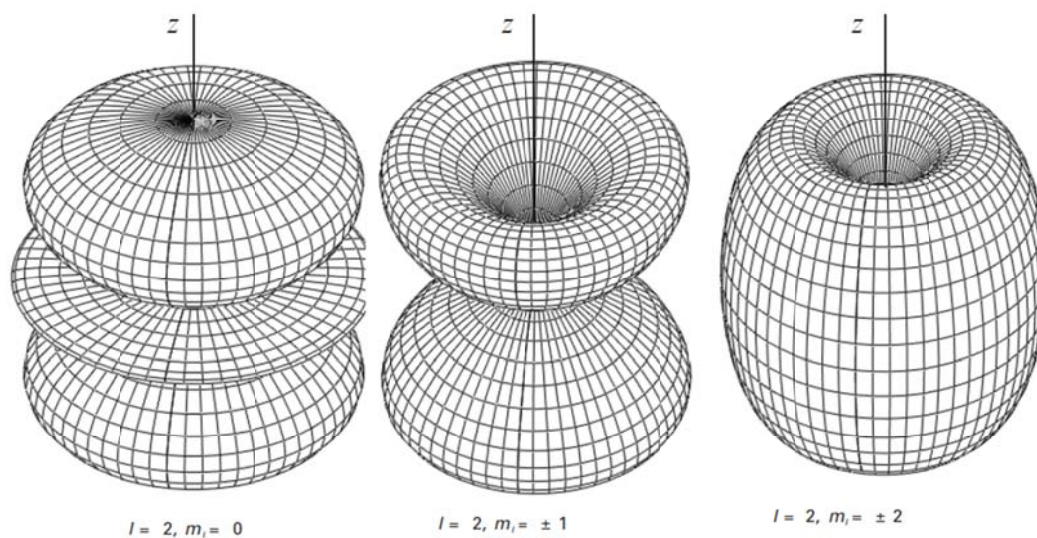


Figure 3.2: A representation of the wavefunctions and the location of the angular nodes for a particle on a sphere with $l = 2$.

Exercise: Draw the corresponding diagrams for $l = 4$.

3.14 Start with eqn 3.45

$$-\frac{\hbar^2}{2m} \nabla^2 \psi = -\frac{\hbar^2}{2m} \left(\frac{1}{r} \frac{\partial^2}{\partial r^2} r + \frac{1}{r^2} \Lambda^2 \right) \psi = E \psi$$

and use the form of the wavefunction in eqn 3.46:

$$\psi(r, \theta, \phi) = R(r)Y(\theta, \phi)$$

This yields (using eqn 3.33):

$$\begin{aligned} -\frac{\hbar^2}{2m} \left(\frac{1}{r} \frac{\partial^2}{\partial r^2} r R Y + \frac{1}{r^2} \Lambda^2 R Y \right) &= -\frac{\hbar^2}{2m} \left(\frac{Y}{r} \frac{d^2}{dr^2} r R + \frac{R}{r^2} \Lambda^2 Y \right) \\ &= -\frac{\hbar^2}{2m} \left(\frac{Y}{r} \frac{d^2}{dr^2} r R + \frac{-R l(l+1)}{r^2} Y \right) \\ &= E R Y \end{aligned}$$

Dividing the last two lines above by Y and multiplying by $-\hbar^2/2m$ results in

$$\frac{1}{r} \frac{d^2(rR)}{dr^2} - \frac{l(l+1)}{r^2} R = -\frac{2m}{\hbar^2} ER$$

which is eqn 3.47a. Writing $k^2 = 2mE/\hbar^2$ and $z = kr$ gives eqn 3.47b.

3.15 Use eqn 3.57 in the form (with $Z = 1$)

$$(1/r)(d^2/dr^2)rR + \{(\mu e^2/2\pi\epsilon_0\hbar^2) - l(l+1)/r^2\}R = -(2\mu E/\hbar^2)R$$

with

$$E = -(\mu e^4/32\pi^2\epsilon_0^2\hbar^2)(1/n^2) \quad [\text{eqn 3.66}]$$

(a) $R_{10} : n = 1, l = 0; E = -(\mu e^4/32\pi^2\epsilon_0^2\hbar^2); l = 0$

$$(1/r)(d^2/dr^2)rR_{10} + (\mu e^2/2\pi\epsilon_0\hbar^2)(1/r)R_{10} = -(2\mu E/\hbar^2)R_{10}$$

Then, because $R_{10} \propto e^{-r/a}$,

$$\begin{aligned} (d^2/dr^2)rR_{10} &= 2R'_{10} + rR''_{10} = -(2/a)R_{10} + (r/a^2)R_{10} \\ &- (2/ar) + (1/a^2) + (\mu e^2/2\pi\epsilon_0\hbar^2)(1/r) = -(2\mu E/\hbar^2) \end{aligned}$$

But $2/a = \mu e^2/2\pi\epsilon_0\hbar^2$; hence $1/a^2 = -2\mu E/\hbar^2$, so $E = -\hbar^2/2\mu a^2$, as required.

(b) $R_{20} \propto (2 - \rho)e^{-\rho/2} = (2 - r/a)e^{-r/2a}$; $E_{2s} = -\frac{1}{4}(\hbar^2/2\mu a^2)$

$$\begin{aligned} (d^2/dr^2)rR_{20} &= 2R'_{20} + rR''_{20} \propto \{-(4/a) + (5r/2a^2) - (r^2/4a^3)\}e^{-r/2a} \\ &- (4/ar) + (5/2a^2) - (r/4a^3) + \underbrace{(\mu e^2/2\pi\epsilon_0\hbar^2)}_{2/a}(1/r)(2 - r/a) \\ &= -(2\mu E/\hbar^2)(2 - r/a) \end{aligned}$$

$$\begin{aligned}
 -(4/ar) + (5/2a^2) - (r/4a^3) + (4/ar) - (2/a^2) &= (1/2a^2) - (r/4a^3) \\
 &= (1/4a^2)(2 - r/a)
 \end{aligned}$$

Hence, $-2\mu E/\hbar^2 = 1/4a^2$, as required.

- (c) $R_{31} \propto (4 - \rho)\rho e^{-\rho/2}$, $\rho = 2r/3a$, $l(l+1) = 2$; then proceed as above, obtaining $-2\mu E/\hbar^2 = 1/9a^2$.

Exercise: Confirm that R_{11} and R_{30} satisfy the wave equation.

3.16 The radial nodes are at the zeros of R_{nl} ; denote them r_0 .

- (a) ψ_{2s} : $R_{2s} = 0$ when $2 - \rho = 0$; $\rho = r/a$.

Hence, $r_0/a = 2$ or $r_0 = 2a = 105.8 \text{ pm}$

- (b) ψ_{3s} : $R_{3s} = 0$ when $6 - 6\rho + \rho^2 = 0$, $\rho = 2r/3a$. The solutions are

$$\rho_0 = 3 \pm \sqrt{3}, \text{ or } r_0 = (3 \pm \sqrt{3})(3a/2) = \underline{1.90a, 7.10a \text{ or } 101 \text{ pm, } 376 \text{ pm}}$$

Exercise: Find (a) the Z -dependence of these node locations, (b) the location of the radial nodes of (i) 2p-orbitals, (ii) 4s-orbitals. [A general point in this connection is that A & S lists the locations of zeros of many functions.]

3.17

$$\begin{aligned}
 \int \psi_{2s}^* \psi_{1s} d\tau &\propto \int_0^\infty R_{20} R_{10} r^2 dr \int_0^\infty (2 - Zr/a) e^{-Zr/2a} e^{-Zr/a} r^2 dr \\
 &\propto \int_0^\infty (2r^2 - Zr^3/a) e^{-3Zr/2a} dr = (2^5 a^3/3^3 Z^3) - (2^5 a^3/3^3 Z^3) = 0
 \end{aligned}$$

Exercise: Confirm that ψ_{2s} and ψ_{3s} are orthogonal.

3.18 Evaluate $|\psi_{ns}|^2 = |Y_{00}|^2 R_{n0}^2 = R_{n0}^2/4\pi$.

$$\psi_{1s}^2(0) = 4(Z/a)^3/4\pi = \underline{(1/\pi)(Z/a)^3} = 2.15 \times 10^{-6} \text{ pm}^{-3} \text{ for hydrogen}$$

$$\psi_{2s}^2(0) = \frac{1}{2}(Z/a)^3/4\pi = \underline{(1/8\pi)(Z/a)^3} = 2.69 \times 10^{-7} \text{ pm}^{-3}$$

$$\psi_{3s}^2(0) = (6/9\sqrt{3})^2(Z/a)^3/4\pi = \underline{(1/27\pi)(Z/a)^3} = 7.96 \times 10^{-8} \text{ pm}^{-3}$$

Exercise: Evaluate the probability density for a 4s-orbital.

3.19

$$\begin{aligned} \langle 1/r^3 \rangle &= \int_0^\infty (1/r^3) R_{21}^2 r^2 dr = \int_0^\infty (1/r) R_{21}^2 dr \\ &= (Z/a)^3 (1/2\sqrt{6})^2 (Z/a)^2 \int_0^\infty (1/r) r^2 e^{-Zr/a} dr = \underline{(1/24)(Z/a)^3} \end{aligned}$$

For a hydrogen atom, this is $2.82 \times 10^{-7} \text{ pm}^{-3}$.

The *general expression* is

$$\langle 1/r^3 \rangle_{nlm_l} = \frac{(Z/a)^3}{n^3 l(l + \frac{1}{2})(l+1)}$$

Exercise: Evaluate **(a)** $\langle 1/r^2 \rangle$ for a 2p_z-orbital, **and (b)** $\langle (1 - 3 \cos^2 \theta)/r^3 \rangle$ for (i) a 2s-orbital, (ii) a 2p_z-orbital.

3.20 $I = hcR$ (i.e. $I = -E_{1s}$)

$$\begin{aligned} I(\text{H}) - I(\text{D}) &= hc(R_{\text{H}} - R_{\text{D}}) = hc(\mu_{\text{H}} - \mu_{\text{D}})e^4/8h^3 \varepsilon_0^2 c \\ &= (\mu_{\text{H}} - \mu_{\text{D}})hcR_\infty/m_e \end{aligned}$$

$$\mu_{\text{H}} = m_e m_p / (m_e + m_p), \quad \mu_{\text{D}} = m_e m_d / (m_e + m_d)$$

$$\begin{aligned}
 (\mu_{\text{H}} - \mu_{\text{D}})/m_e &= \frac{m_p}{m_e + m_p} - \frac{m_d}{m_e + m_d} \\
 &= \frac{m_e(m_p - m_d)}{(m_e + m_p)(m_e + m_d)} = \frac{m_e(m_{\text{H}} - m_{\text{D}})}{m_{\text{H}}m_{\text{D}}}
 \end{aligned}$$

$$m_e = 9.109\,38 \times 10^{-31} \text{ kg}, m_{\text{H}} = 1.6735 \times 10^{-27} \text{ kg}, m_{\text{D}} = 3.3443 \times 10^{-27} \text{ kg};$$

$$(\mu_{\text{H}} - \mu_{\text{D}})/m_e = -2.7195 \times 10^{-4}$$

Consequently,

$$\begin{aligned}
 I(\text{H}) - I(\text{D}) &= -(2.7195 \times 10^{-4}) \times (2.1799 \times 10^{-18} \text{ J}) \\
 &= -5.9282 \times 10^{-22} \text{ J} = \underline{-0.357 \text{ kJ mol}^{-1}} (-3.70 \text{ meV})
 \end{aligned}$$

The experimental values are $109\,678.758 \text{ cm}^{-1}$ and $109\,708.596 \text{ cm}^{-1}$, so

$$\{I(\text{H}) - I(\text{D})\}/\text{cm}^{-1} = -29.838 \text{ cm}^{-1} (-0.357 \text{ kJ mol}^{-1})$$

Exercise: Evaluate the ionization energy of positronium on the basis of the ionization energy of ^1H .

3.21 For a given value of l there are $2l + 1$ values of m_l . For a given n there are n values of l .

Hence, the degeneracy g is

$$g = \sum_{l=0}^{n-1} (2l + 1) = n(n - 1) + n = \underline{n^2}$$

Exercise: Calculate the average value of m_l^2 for an atom in a state with principal quantum number equal to n but with l, m_l unspecified.

Problems

3.1 Write $x = r \cos \phi$, $y = r \sin \phi$, $r = (x^2 + y^2)^{1/2}$, $\phi = \arctan(y/x)$. Then

$$\frac{\partial f}{\partial x} = \frac{\partial f}{\partial r} \frac{\partial r}{\partial x} + \frac{\partial f}{\partial \phi} \frac{\partial \phi}{\partial x} = \left(\cos \phi \frac{\partial}{\partial r} - \frac{\sin \phi}{r} \frac{\partial}{\partial \phi} \right) f$$

or

$$\frac{\partial}{\partial x} = \cos \phi \frac{\partial}{\partial r} - \frac{\sin \phi}{r} \frac{\partial}{\partial \phi}$$

Similarly,

$$\frac{\partial}{\partial y} = \sin \phi \frac{\partial}{\partial r} + \frac{\cos \phi}{r} \frac{\partial}{\partial \phi}$$

Therefore,

$$\begin{aligned} \frac{\partial^2 f}{\partial x^2} &= \frac{\partial}{\partial x} \left(\frac{\partial f}{\partial x} \right) = \left(\cos \phi \frac{\partial}{\partial r} - \frac{\sin \phi}{r} \frac{\partial}{\partial \phi} \right) \left(\cos \phi \frac{\partial}{\partial r} - \frac{\sin \phi}{r} \frac{\partial}{\partial \phi} \right) f \\ &= \left\{ \cos^2 \phi \frac{\partial^2}{\partial r^2} - \sin \phi \cos \phi \frac{\partial}{\partial r} \left(\frac{1}{r} \frac{\partial}{\partial \phi} \right) \right. \\ &\quad \left. - \frac{\sin \phi}{r} \frac{\partial}{\partial \phi} \left(\cos \phi \frac{\partial}{\partial r} \right) + \frac{\sin \phi}{r^2} \frac{\partial}{\partial \phi} \left(\sin \phi \frac{\partial}{\partial \phi} \right) \right\} f \end{aligned}$$

That is,

$$\begin{aligned} \frac{\partial^2}{\partial x^2} &= \cos^2 \phi \frac{\partial^2}{\partial r^2} - \frac{2 \sin \phi \cos \phi}{r} \frac{\partial^2}{\partial r \partial \phi} + \frac{\sin^2 \phi}{r^2} \frac{\partial^2}{\partial \phi^2} \\ &\quad + \frac{\sin^2 \phi}{r} \frac{\partial}{\partial r} + \frac{2 \sin \phi \cos \phi}{r^2} \frac{\partial}{\partial \phi} \end{aligned}$$

Similarly,

$$\frac{\partial^2}{\partial y^2} = \sin^2 \phi \frac{\partial^2}{\partial r^2} + \frac{2 \sin \phi \cos \phi}{r} \frac{\partial^2}{\partial r \partial \phi} + \frac{\cos^2 \phi}{r^2} \frac{\partial^2}{\partial \phi^2} + \frac{\cos^2 \phi}{r} \frac{\partial}{\partial r} - \frac{2 \sin \phi \cos \phi}{r^2} \frac{\partial}{\partial \phi}$$

It then follows that

$$\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} = \frac{\partial^2}{\partial r^2} + \frac{1}{r} \frac{\partial}{\partial r} + \frac{1}{r^2} \frac{\partial^2}{\partial \phi^2}$$

as in eqn 3.2.

Exercise: Derive an expression for ∇^2 in cylindrical polar coordinates, $x = r \cos \phi$, $y = r \sin \phi$, z .

3.4

$$\langle r \rangle = \int_0^\infty r P(r) dr = \int_0^\infty R^2 r^3 dr \quad [\text{eqn 3.69}]$$

$$\langle 1/r \rangle = \int_0^\infty r^{-1} P(r) dr = \int_0^\infty R^2 r dr \quad [\text{eqn 3.69}]$$

As in Problem 3.3,

(a)

$$\langle r \rangle = \left(\frac{Z^3}{243 a_0^3} \right) \left(\frac{3 a_0}{2Z} \right)^4 \int_0^\infty (6 - 6\rho + \rho)^2 \rho^3 e^{-\rho} d\rho = \frac{27 a_0}{2Z}$$

$$\langle 1/r \rangle = \left(\frac{Z^3}{243 a_0^3} \right) \left(\frac{3 a_0}{2Z} \right)^2 \int_0^\infty (6 - 6\rho + \rho^2)^2 \rho e^{-\rho} d\rho = \frac{Z}{9 a_0}$$

(b)

$$\langle r \rangle = \left(\frac{Z^3}{81 \times 6a_0^3} \right) \left(\frac{3a_0}{2Z} \right)^4 \int_0^\infty (4-\rho)^2 \rho^5 e^{-\rho} d\rho = \frac{25a_0}{2Z}$$

$$\langle 1/r \rangle = \left(\frac{Z^3}{81 \times 6a_0^3} \right) \left(\frac{3a_0}{2Z} \right)^2 \int_0^\infty (4-\rho)^2 \rho^3 e^{-\rho} d\rho = \frac{Z}{9a_0}$$

We have used the integrals

$$\int_0^\infty (6-6x+x^2)^2 x^3 e^{-x} dx = 648$$

$$\int_0^\infty (6-6x+x^2)^2 x e^{-x} dx = 12$$

$$\int_0^\infty (4-x)^2 x^5 e^{-x} dx = 1200$$

$$\int_0^\infty (4-x)^2 x^3 e^{-x} dx = 24$$

as obtained by using the symbolic integration procedure in mathematical software.

Exercise: Evaluate $\langle 1/r^3 \rangle$ for each orbital.

3.7 Use available mathematical software to find zeroes of the Bessel functions; in particular find values of z such that $J(z) = 0$. With z identified as ka (see eqn 3.25), the energies can be expressed in terms of z as

$$E = \frac{k^2 \hbar^2}{2m} = \frac{z^2 \hbar^2}{2ma^2}$$

3.10 (a) The moment of inertia of a sphere is $I = \frac{2}{5} MR^2$ [Problem 10.1]; therefore, on

writing this value as Mr^2 , we see that $r = (2/5)^{1/2} R$.

- (b) Consider rotation perpendicular to the axis. The mass of a disc of thickness dx , radius R is $\pi\rho R^2 dx$ where ρ is the mass density of the disc. Therefore,

$$I = \int_{-l/2}^{l/2} \pi\rho R^2 x^2 dx = \frac{1}{12} \pi\rho R^2 l^3$$

The mass of the cylinder is $M = \pi\rho R^2 l$; therefore $I = \frac{1}{12} Ml^2$.

Setting this value equal to $M r^2$ gives $r = l/(12)^{1/2}$

- 3.13** The wavefunction for a particle in a spherical cavity is given by

$$\psi = Nj(r)Y(\theta, \phi)$$

The ground-state wavefunction is therefore given by $Nj_0 Y_{0,0}$. (i) Proceeding as in Problem 3.8, (ii) using eqn 3.48 for j_0 with $k = \pi/a$ (Table 3.3) and (iii) recognizing that the volume element contains a factor of $r^2 dr$, we write the probability for finding the particle within a sphere of radius $a/2$ as:

$$P = \frac{\int_0^{a/2} \frac{\sin^2 kr}{(kr)^2} r^2 dr}{\int_0^a \frac{\sin^2 kr}{(kr)^2} r^2 dr} = \frac{\int_0^{a/2} \sin^2 kr dr}{\int_0^a \sin^2 kr dr} = \frac{\int_0^{a/2} \sin^2\left(\frac{\pi r}{a}\right) dr}{\int_0^a \sin^2\left(\frac{\pi r}{a}\right) dr}$$

Using mathematical software or standard integration tables yields $P = 1/2$.

- 3.16** Refer to Fig. 3.5. The rotation (a)→(b) corresponds to 3p becoming 3d, the rotation (b)→(c) corresponds to 3d becoming 3s.

Exercise: Identify the patterning of the ball that would account for the degeneracy of two-dimensional f-orbitals.

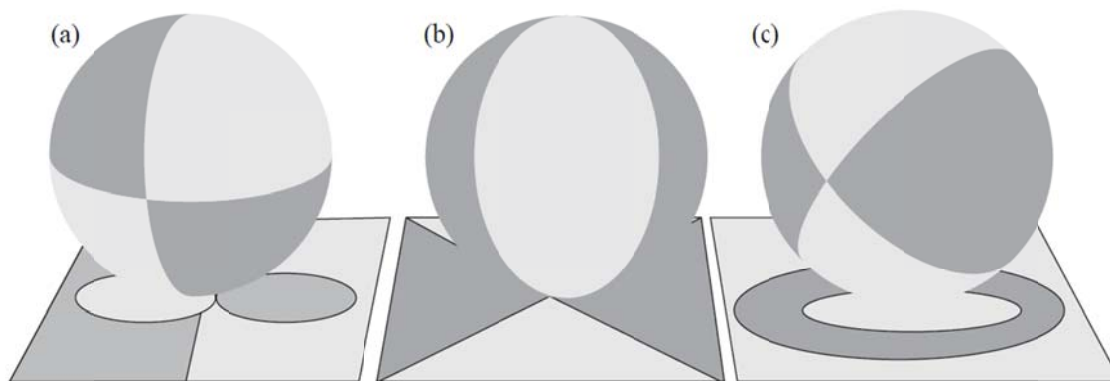


Figure 3.5: The projections of a patterned sphere on a plane (the projection stems from the North Pole).